

pA Physics in STAR

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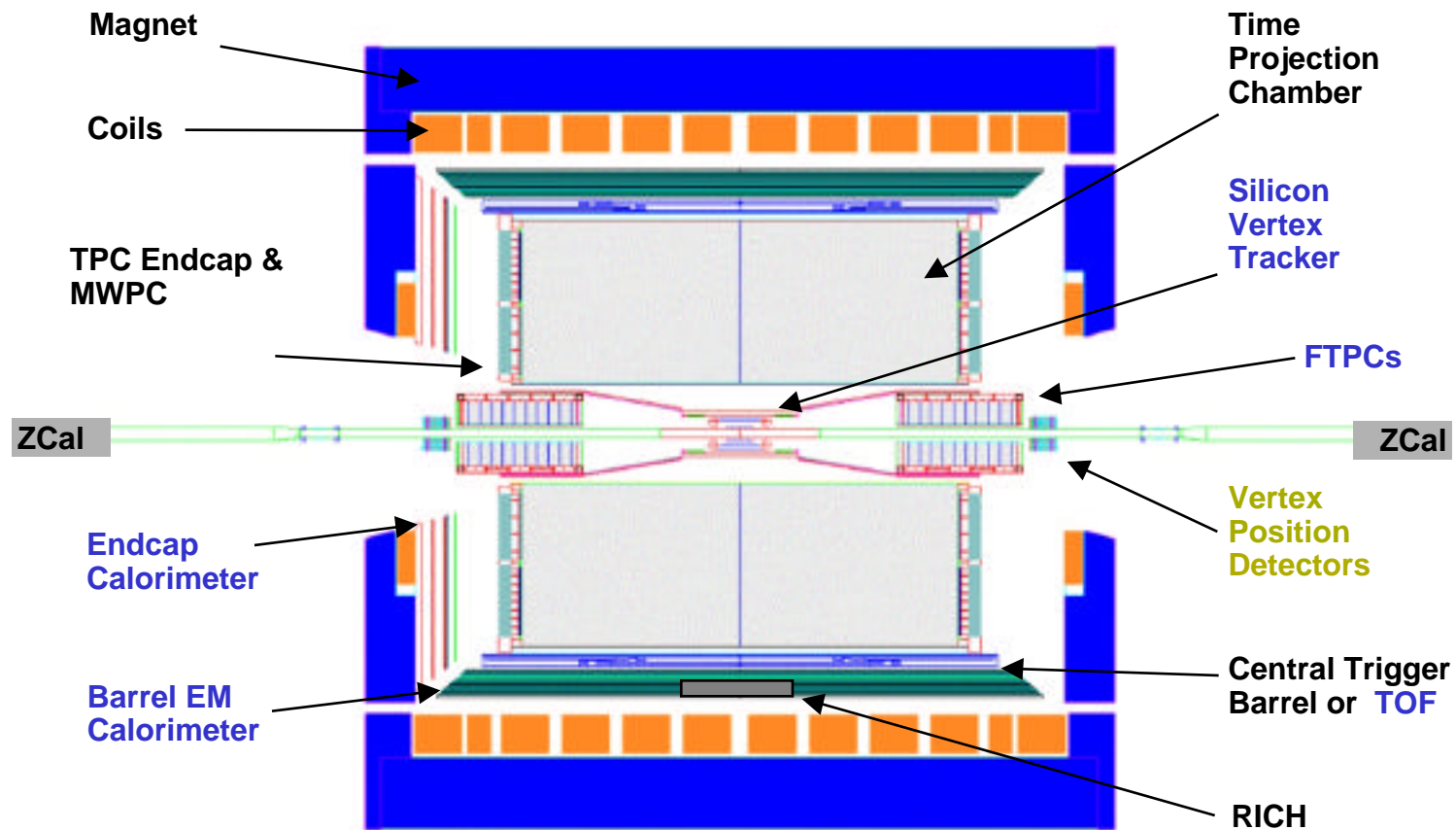
- STAR detector and pA-related observables
- pA and AA physics
- Gluon Distributions in Nuclei
- Event Characterization in pA
- STAR upgrades relevant to pA
- Summary

Preliminary remarks

- pA physics is often mentioned as part of STAR experimental program
- Some individual pA studies (G. Odyniec, H. Huang) but until now no formal framework within collaboration
- pA at RHIC is of broad interest:
 - Fundamental study of QCD
 - Isolation of nuclear effects to interpret AA data
- I will present some thoughts on pA in STAR
 - Not yet a proposal for an experimental program
 - Many open questions

Aim of this talk: set scale of pA program, identify some important issues for STAR to address

The STAR Detector



RICH replaced by final EMC modules in '03

STAR Institutions

U.S. Labs:

**Argonne, Berkeley, and
Brookhaven National Labs**

U.S. Universities:

**Arkansas, UC Berkeley, UC Davis,
UCLA, Carnegie Mellon, Creighton,
Indiana, Kent State, MSU, CCNY,
Ohio State, Penn State, Purdue,
Rice, Texas A&M, UT Austin,
Washington, Wayne State, Yale**

Brazil:

Universidade de Sao Paolo

China:

IHEP - Beijing, IPP - Wuhan

England:

University of Birmingham

France:

**Institut de Recherches Subatomiques
Strasbourg, SUBATECH - Nantes**

Germany:

**Max Planck Institute – Munich
University of Frankfurt**

Poland:

**Warsaw University
Warsaw University of Technology**

Russia:

**MEPHI – Moscow, LPP/LHE JINR –
Dubna, IHEP - Protvino**

Offline Reconstruction (Au+Au at 130 GeV)

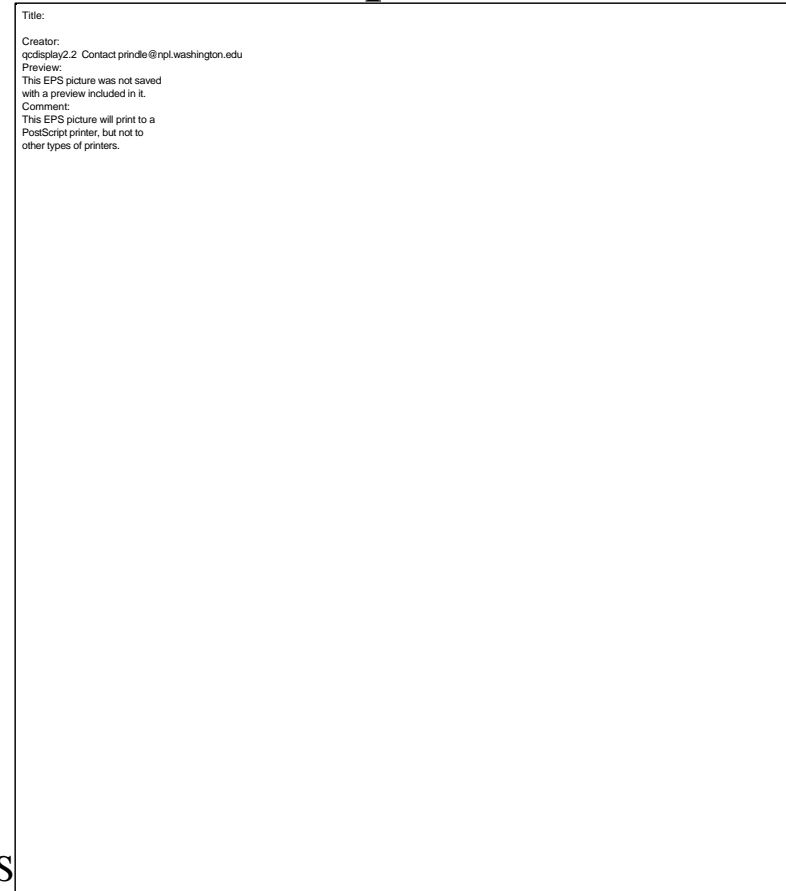
(modest multiplicity shown so details are visible)

All hits on tracks:

- red = low pT
- violet = high pT



Hits not assigned to tracks:
mostly noise and low
momentum spirals



s in S

Barrel EM Calorimeter

- Coverage: $-1 < \eta < 1$
- 4800 lead-scintillator projective towers ($18 X_0$)
- each tower $x = 0.05 \times 0.05$
- resolution $E/E = 16\%/\sqrt{E}$
- preshower (factor 4 hadron rejection for electrons)
- gas-filled shower max at $5X_0$:
 - 36000 channels
 - resolve photons from π^0 at 25 GeV

Endcap EM Calorimeter

- Coverage: $1.07 < \eta < 2$
- 720 lead-scintillator projective towers ($21 X_0$, sampling fraction 6.6%)
- towers $x = 0.05 - 0.10 \times 0.10$
- resolution $E/E \sim 14\%/\sqrt{E}$
- preshower (factor 4 hadron rejection for electrons)
- scintillator shower max at $5X_0$:
 - 7200 channels
 - rejects 80% of π^0 s with 80% photon efficiency at 30 GeV

General Remark on Calorimetry

- Calorimeters are **essential** for triggering on high pT jets, hadrons (π^0), photons and electrons
- Important fact: installation schedule:
 - Barrel: 25% per year from FY01 to early FY04
 - Endcap: 50% FY02, 50% FY03
- High p_T component of pA in STAR requires calorimeter
Extended pA running only makes sense starting in '03

Main STAR Observables for pA

- Charged particles to high p_T (~ 15 GeV)
- π^0 , γ to high p_T
- Jets: inclusive, coincidence
- Identified hadrons
 - $\pi/K/p/\dots$ via dE/dx in TPC ($p < 1$ GeV)
 - topological PID (K^+ , K^- , K_s^0 , Λ , $\bar{\Lambda}$, $\bar{\Lambda}$, \dots)
 - in some cases, $p_T \sim$ several GeV
 - Resonances
- J/ψ , e^+e^- , D , K

Physics topic 1: Jets in AA and pA

- Energy loss dE/dx of high energy partons in medium is sensitive to energy density
- dE/dx via gluon brehmsstrahlung: jet fragmentation profile changes (softens)
- Energy loss in AA measured in STAR via:
 - High p_T hadron inclusive spectra+coincidences (π^0 , charged hadrons)
 - Flavor tagging (leading K^+/K^- , $\Lambda/\bar{\Lambda}$, p/\bar{p})
quark/gluon differential energy loss
 - Gamma/jet coincidence: photon is non-interacting “standard candle”
 - (“flow” of high p_T hadrons wrt reaction plane)

Jets in AA and pA (cont'd)

- Initial state effects also influence jet/hadron spectrum:
 - shadowing
 - generation of k_T via multiple soft interactions
- X.N. Wang et al.:
 - pp and pA measurements may isolate initial state effects
 - pp and pA data are essential to interpret AA data

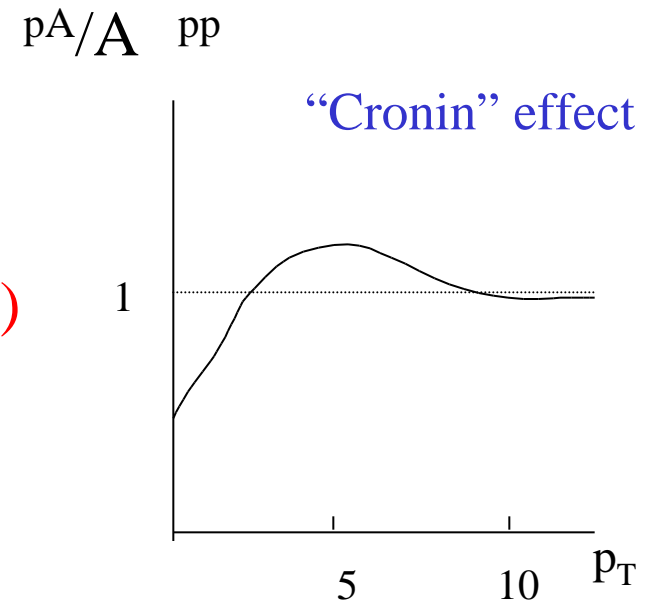
Physics topic 2: Inclusive Hadron Spectrum

AA: p_T distribution of hadrons
influenced by multiple
scattering and radial flow

pA: multiple scattering only (+shadowing)

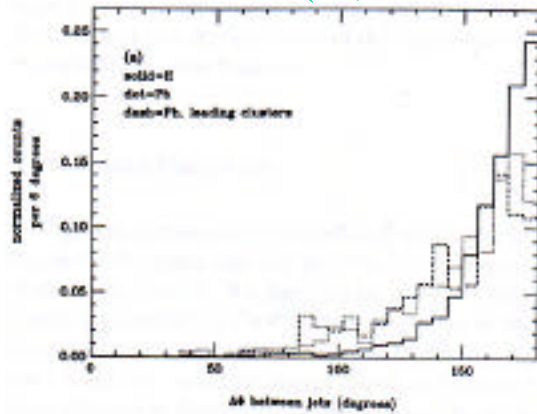
Rough estimate of hadron yield limits
(~100 counts/GeV, 10 weeks of p+Au):
pions to $p_T > 15$ GeV/c; Kaons to $p_T > 12$
GeV/c; / bar to $p_T \sim 10$ GeV/c

Topological PID in TPC: how high in
 p_T ?



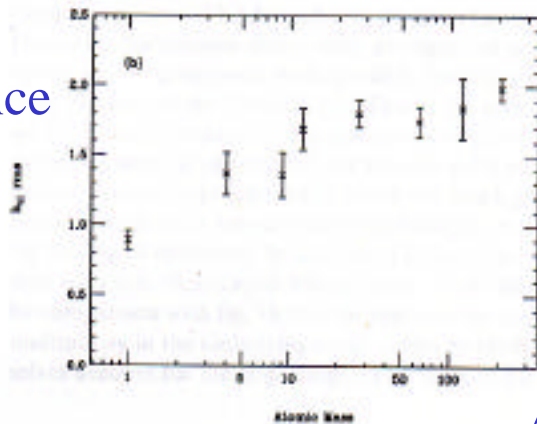
Physics topic 3: Multiple Scattering via Dijets in pA

Angular correlations (H,
Pb)



(jet-jet)

k_T mass
dependence



A

- E609: dijet acoplanarity due to multiple scattering
- Larger effects in Pb than H in
 - fitted k_T
 - acoplanarity
- High quality measurements look possible in STAR (but not yet studied):
 - Robust dijet yields
 - Good dijet acceptance, energy resolution

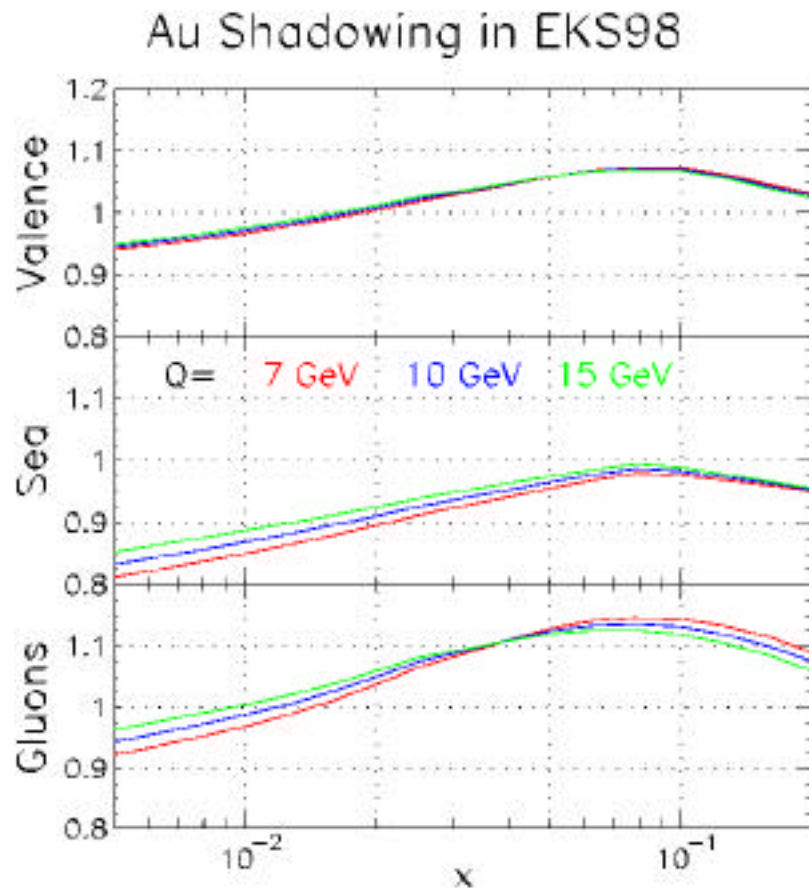
Physics topic 4: Charm in AA and pA

- AA collisions: J/ψ suppression is important signature of deconfinement (Debye screening)
- STAR will measure:
 - $J/\psi \rightarrow e^+e^-$ (low x_F)
 - $D \rightarrow K$
- Fermilab E866: J/ψ measured in pA
 - suppression at all x_F (~ 0.95 @ $x_F \sim 0$)
 - p_T dependence consistent with multiple scattering
- Charm measurement in pA is essential to interpret AA data!
- A-dependence is important
- STAR: need to study J/ψ capabilities specific to pA

Physics topic 5: Soft hadrons in pA and AA

- pA is essential reference for AA soft physics: baryon stopping, strangeness enhancement, ...
- No additional experimental preparation needed for pA (except trigger?)
- Modest statistics needed except for rare probes (e.g.)
 - No event rate estimates done yet to set scale of program
 - Hard probes more demanding for beam time soft physics will mostly be done along the way
- Systematics: A-dependence
- Energy scan at modest statistics?
- Event characterization?
- Minbias Triggering?

Physics topic 6: Measurement of Gluon Distribution in Nuclei



(Carl Gagliardi/E866)

- Effects are small ($\sim 10\%$) requires precision measurement
- **STAR: Compton process**
 $q+g \rightarrow q+\gamma$
 - Dominates direct photon xsection (90%)
- Low x asymmetric collision EMC Endcap is essential

Calculation of Statistical Precision of Direct Photon Measurement

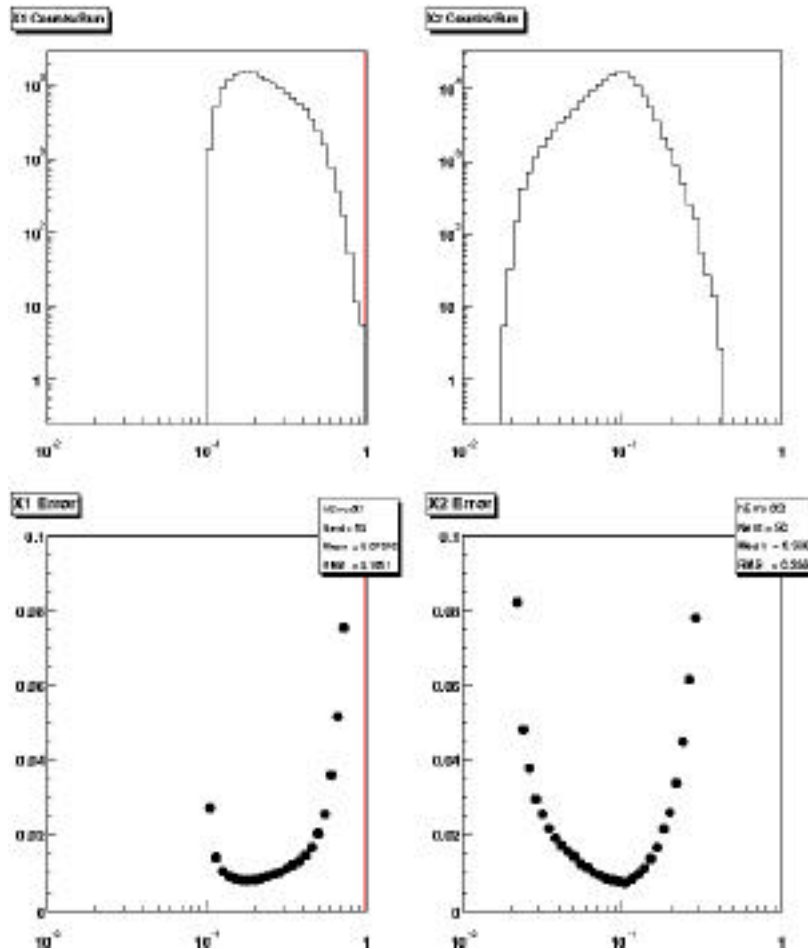
- Hijing, p+Au @ $\sqrt{s}=200$ GeV, $p_T > 10$ GeV
- Integrated luminosity:
 - Sam Aronson: best guess is $\sim 5 \times 10^{28} / \text{cm}^2 / \text{sec}$
 - Assume 10 week run (probably generous), 75% duty factor **Integrated Luminosity $\sim 2 \times 10^5 / \mu\text{b}$**
- STAR acceptance: Parton/ kinematics, no jet finding
 - : $-1 < < 2$, jet: $-0.5 < < 1.5$
- Measure p_T , , x_1, x_2
- Assume $x_{\text{gluon}} = x_2 (< x_1)$

Why $p_T > 10$ GeV?

- In principle, small p_T interesting to access small x_{gluon}
- Instrumental issue: σ too small at lower p_T
 - Result from pp (spin) studies, should be same in pA
- Theoretical issues:
 - Large fragmentation photon yield at smaller p_T
 - E706 effect: NLO not sufficient at $p_T < 10$ GeV (but what about collider kinematics – lower x_T ?)
 - Additional k_T due to multiple scattering in nuclei: bigger relative effect at smaller p_T

We restrict our discussion to $p_T > 10$ GeV

Statistical Precision for x1, x2 (10 week run, full EMC coverage)



- Upper panels: counts per bin
- Lower panels: relative statistical error

5% @ $x_2 \sim .02$

( sharply falling spectrum)

- 10 weeks needed for high quality measurement
- '02 run: partial EMC coverage

5% @ $x_2 \sim .04$

Summary of Gluon Distribution Measurement

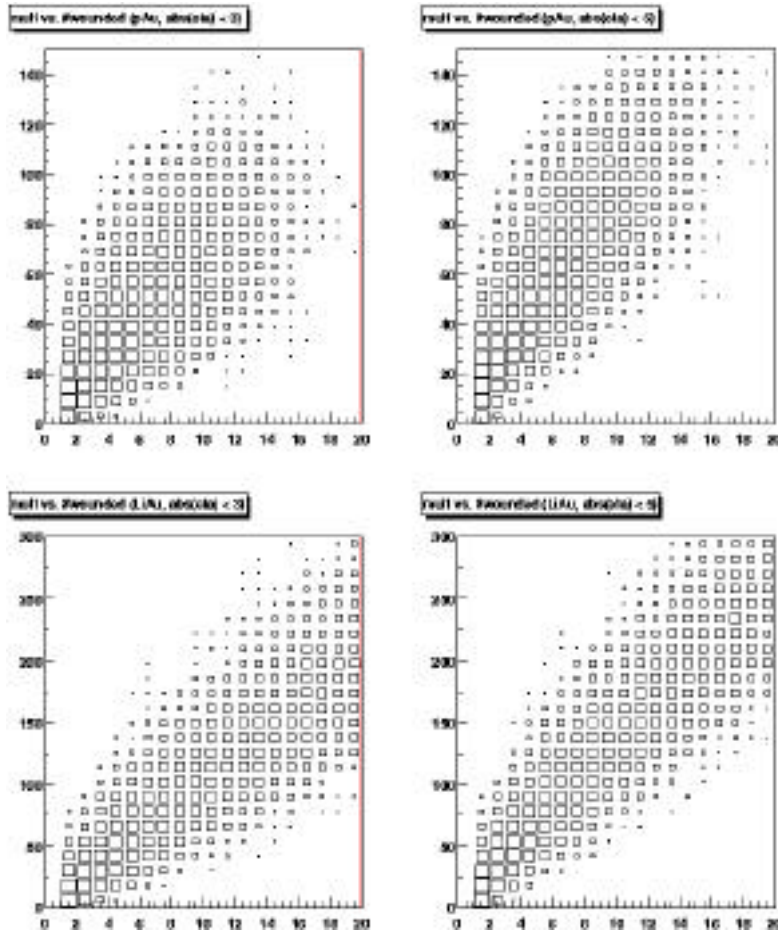
- From statistical point of view looks doable with ~10 weeks running per year per A
- Study A-dependence (3 masses?) multiyear program
- Open experimental and theoretical issues:
 - $\sqrt{s_{NN}}$
 - Fragmentation photon yields and isolation cuts
 - resolutions
 - Luminosity monitoring (Sebastian's talk):
 - Van der Meer scan? Other ideas?

Event characterization in pA

- Tantalizing prospect: vary thickness of target via tagging of collision geometry (# excited target nucleons)
- Famous problem in fixed target: measure “grey” protons (Brian Cole’s talk)
 - Very tough to measure in collider mode
- Alternative: charged multiplicity in e.g. $-3 < \eta < 3$
 - Correlation with # excited target nucleons not bad (see next slide)
- Another alternative: multiplicity in e.g. Li+Au
 - Advantage: correlation w/ geometry significantly better
 - Disadvantage: “dirty” probe for high precision work
 - But how dirty? Needs some study

Characterization of Geometry via Charged Multiplicity

z axis is logarithmic



- Hijing minbias @200 GeV for p+Au, Li+Au
- Charged multiplicity within $|\eta| < 3$ and $|\eta| < 5$
- Multiplicity vs # wounded “target” nucleons
- correlation exists in p+Au good enough?
- correlation significantly sharper in Li+Au (note log scale in z)

STAR Upgrades relevant to pA

- High precision vertex tracker for D mesons hadrons
- More forward coverage for yet lower x_{gluon} ??
 - Low x measured by very asymmetric collisions EMC Endcap
 - Consider even more forward calorimeter to get lower x_{gluon} (e.g. $3 < x < 5$ stacked around the beam pipe)
 - $p_T > 10$ GeV: unfortunately, cross sections are tiny
no rate
 - $p_T > 5$ GeV: plenty of rate, but unfavourable ϕ ,
interpretational problems

Forget it!

Summary of pA Physics Topics

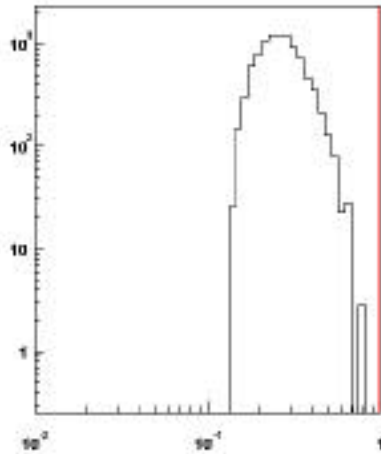
- Jets and high p_T hadrons:
 - pA essential to disentangle mult scattering effects
 - STAR has many observables, robust capabilities
- Charm:
 - Essential to measure J/ψ absorption in pA
 - STAR measures well at small x_F
- Soft hadrons:
 - pA supplies essential reference, STAR has many channels
 - Energy scan, event characterization, triggering,...
- Gluon distributions in nuclei:
 - Probably doable statistically
 - Many experimental and theoretical issues

Conclusions

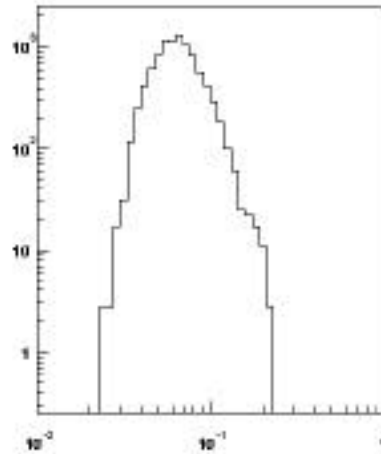
- pA is a major program at RHIC
 - Important to measure A-dependence
 - Several weeks of running per year over several years
- General experimental issues: trigger, luminosity monitoring,...
- STAR: extended pA running should begin '03 (driven by calorimeter installation schedule)

Statistical Precision for x1,x2 with “year ’02” EMC coverage

X1 Count/Run

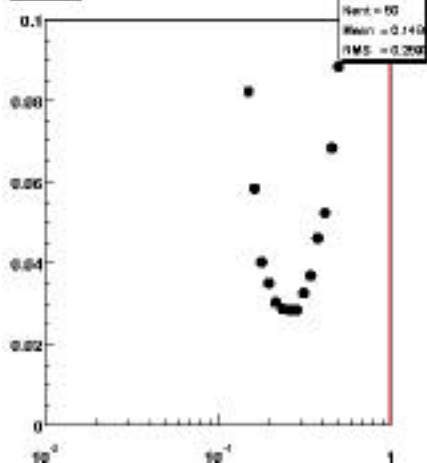


X2 Count/Run

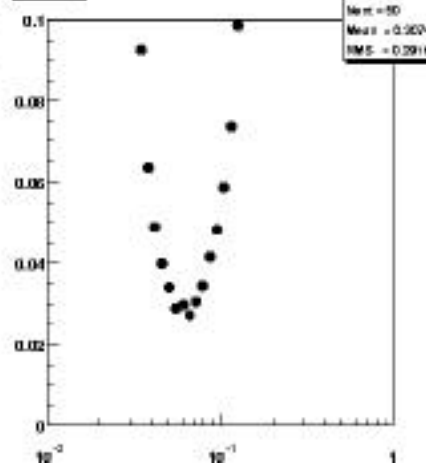


- Barrel:
 - $0 < < 1$
 - $0 < < 2$

X1 Error



X2 Error



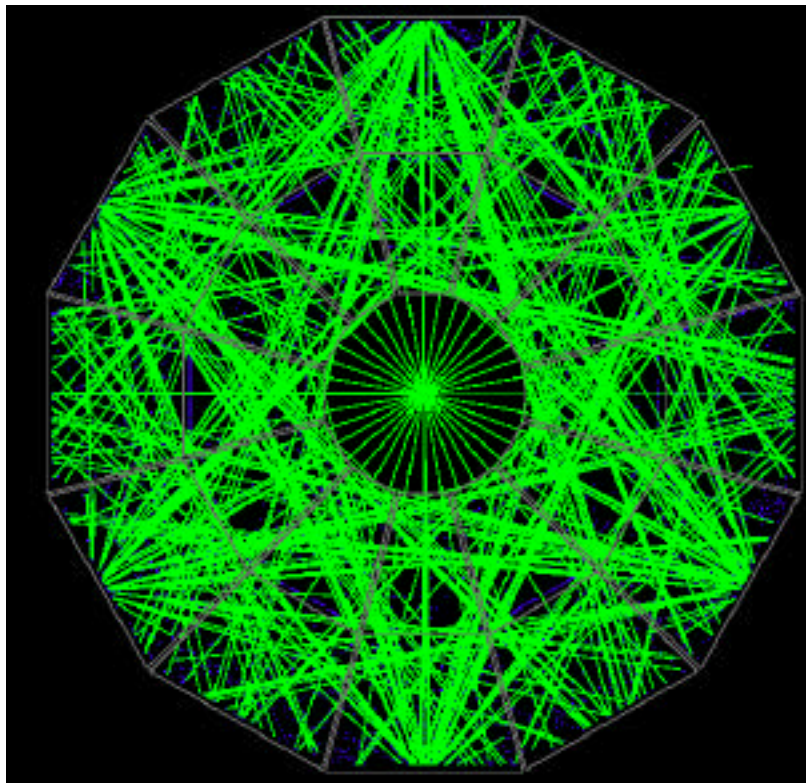
- Endcap:
 - $1 < < 2$
 - $0 < <$

BNL, Oct 28,2000

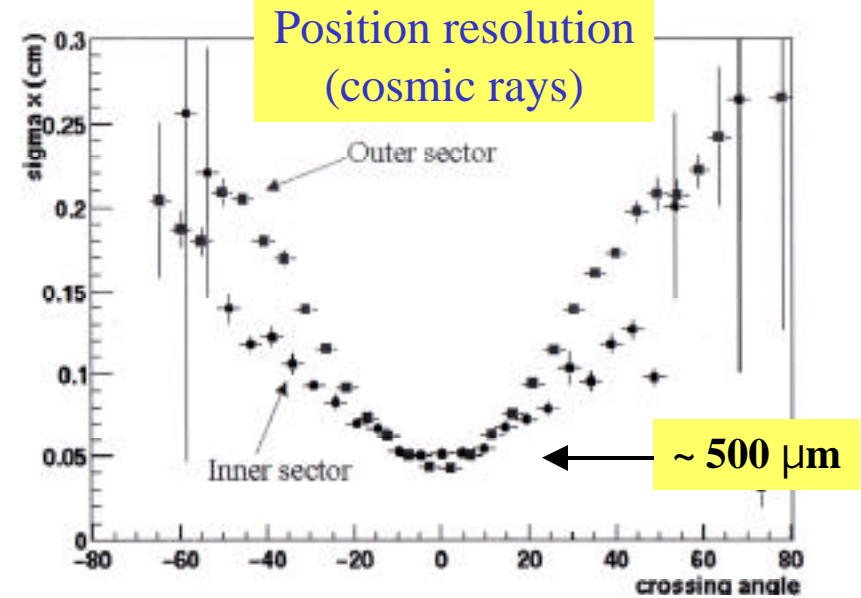
pA Physics in STAR

Track Reconstruction

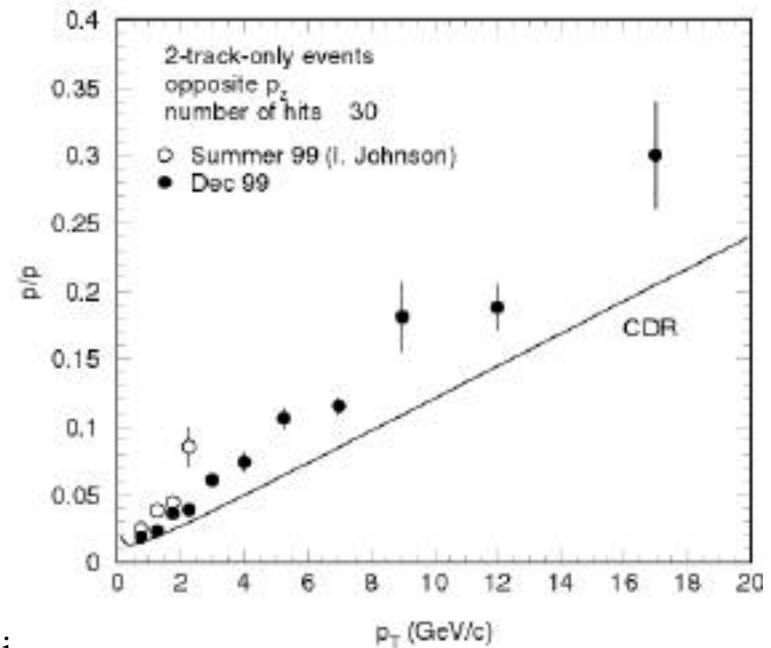
Reconstructed Laser Event



BNL, Oct 28, 2000



Momentum resolution dp/p (cosmics)



pA Physics in STAR